Observing Filaments Between Galaxy Clusters with NenuFAR How to Probe Cosmic Magnetic Fields

NenuFAR workshop presentation Etienne Bonnassieux, postdoctoral research fellow at UniBo

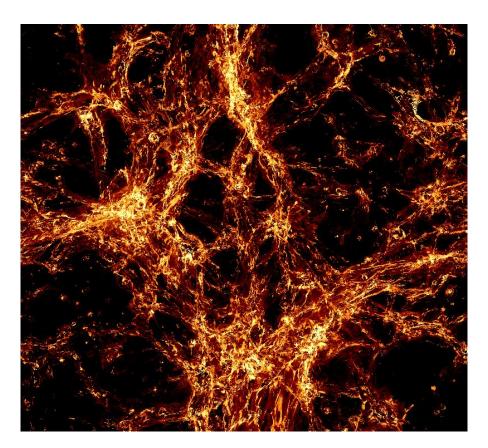
Collaboration with A. Bonafede, F. Vazza, C. Ferrari, F. de Gasperin, C. Tasse, ...

Outline

- Science Goals

- Why NenuFAR?

- Why us?



To study cosmic magnetism:

- Magnetic field in clusters dominated by ICM evolution
- To understand primordial magnetic field, need to go outside clusters

BUT

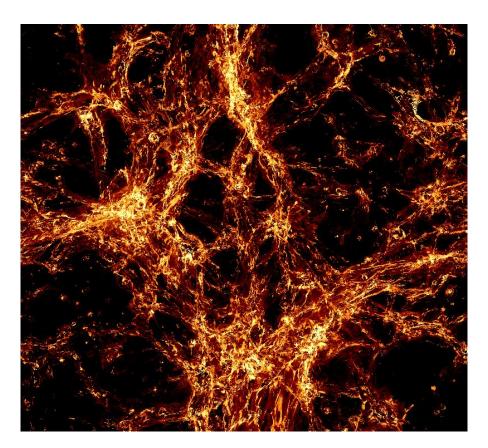
 If there are no relativistic electrons to decay in a forest, what can astronomers know about its magnetic field?...

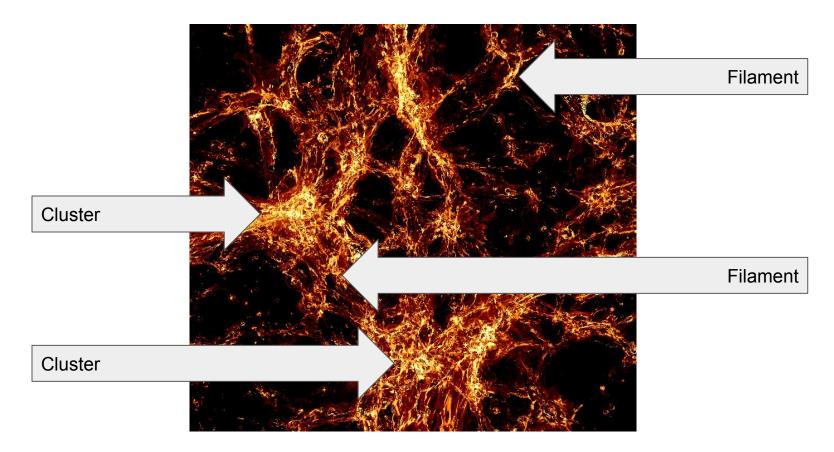
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Filaments: best environment for both presence of emitting particles and

unmarred cosmic magnetic field







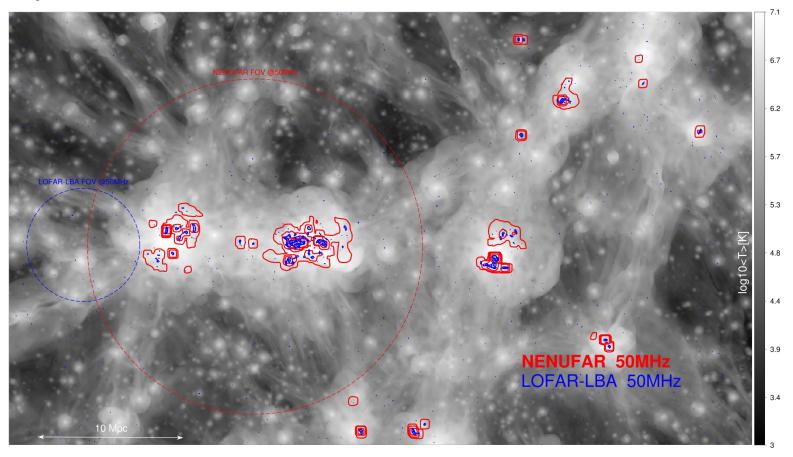
Observational constraints for filament observations:

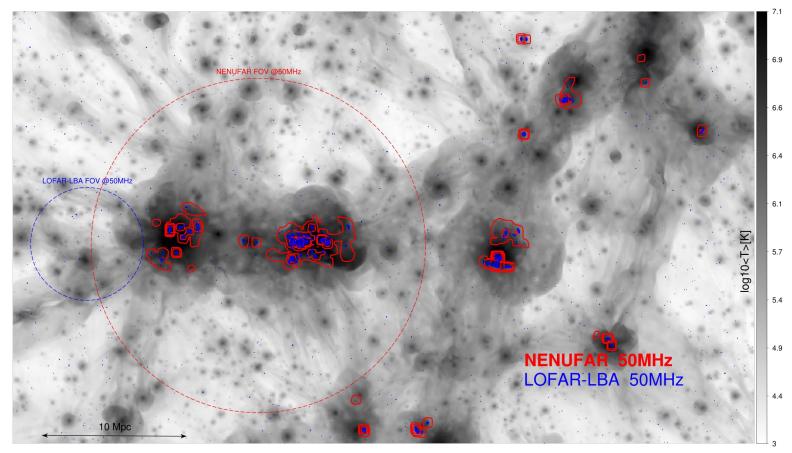
- Physics of cosmic magnetism traced by large-scale, diffuse structures
- Emission mechanism of said structures more effective at lower frequencies
- Presence of foreground/embedded point sources imposes angular resolution constraints to improve confusion limits

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NenuFAR standalone in imager mode + supersynthesis





The Team

KP currently has the following participants:





Etienne Bonnassieux, Annalisa Bonafede, Franco Vazza, Chiara Ferrari,

Francesco de Gasparin, Cyril Tasse, ...









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And maybe...you? ;)